

Astronomy STANDARD LEVEL Paper 2

Friday 29 April 2011 (morning)

90 minutes

Candidate session number									
0	0								

INSTRUCTIONS TO CANDIDATES

- Write your session number in the boxes above.
- Do not open this examination paper until instructed to do sq.
- Section A: Answer ALL of Section A in the spaces provided.
- Section B: Answer ONE question from Section B in the spaces provided.
- There are two sections answer ALL questions in all four sections.

ADDITIONAL INSTRUCTIONS

- Calculators are allowed.
- A 1-page *Information Sheet* is provided for this examination.

Marking Grid							
Section A	Marks						
1							
2							
3							
4							
Section B	Marks						
5							
6							
Total / 60							

The following information may be useful.

$$1 \text{ AU} = 1.496 \times 10^{11} \text{ m}$$

1 parsec =
$$206265 \text{ AU} = 3.09 \times 10^{16} \text{ m} = 3.26 \text{ light years}$$

1 light year =
$$0.307$$
 parsecs = 9.47×10^{15} m

$$c = 3 \times 10^8 \text{ m s}^{-1}$$

$$G = 6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$$

$$\sigma = 5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$$

$$L_{\odot}\approx 3.84 \; \mathrm{x} \; 10^{26} \; \mathrm{W}$$

$$T_{\odot} \approx 5770 \text{ K}$$

$$M_{\odot} \approx 1.99 \times 10^{30} \text{ kg}$$

$$R_{\odot} \approx 6.96 \text{ x } 10^8 \text{ m}$$

$$H_o \approx 72 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

$$z = \frac{H_0}{c} d$$

$$F = \frac{L}{4\pi d^2}$$

$$c = f \lambda$$

$$\lambda_{max} \, = \, \frac{2.90 \, x \, 10^{\text{-}3}}{T}$$

$$PE = -\frac{GMm}{r}$$

$$E = mc^2$$

$$d = \frac{1}{\phi}$$

$$\frac{b_1}{b_2} = 2.5^{(m_2 - m_1)}$$

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Section A

Answer ALL questions You are advised to spend no more than 60 minutes on this section

1. This question is about Cepheid variables.

A Cepheid variable is a giant star and they constitute about two thirds of all variable stars. Figure 1 whows the variation of the apparent magnitude as a function of time.

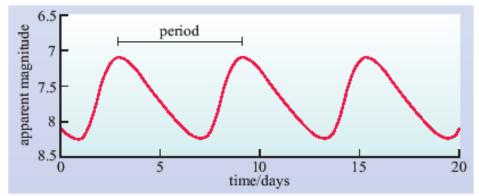


Figure 1. The variation of apparent magnitude for a Cepheid variable

rigure 1. The	variation of apparent magnitude for a cephera variable.	
a) Briefly explain what causes the	ne star's luminosity to vary in the way showp in Figure 1.	
		[2 marks]
b) Using Figure 1, estimate the bthe Cepheid.	orightest apparent magnitude, faintest apparent magnitude a	nd period for
Brightest apparent magnitude :		
Faintest apparent magnitude :		
Period of variation :	days	[2 marks]

c) Using your values for apparent magnitude, show that the ratio of maximum to minimum brightness for the Cepheid is approximately 3.2.
[2 marks]
d) Figure 2 shows how the Cepheid is used as a standard candle. This Figure indicates how the luminosity and period are related.
Figure 2. Luminosity – Period relationship for a Cepheid variable. Use Figure 2 to find the luminosity of the Cepheid – ensure that you make it clear how you have come to your answer and express your answer in multiples of the solar luminosity.
[2 marks]
e) If the luminosity of a star is known and the flux density is measured, the distance to the star can be found.
Given that the flux density of the Cepheid is $3.2 \times 10^{-6} \text{ W/m}^2$, show that the distance to the Cepheid is approximately $1.0 \times 10^{17} \text{m}$.
[2 marks]

2. This question is about comets.

Comets typically consists of a compact nucleus, an atmosphere of gas and dust (the *coma*) and a tail directed away from the Sun. The comet is made of lumps of ice mixed with silicate particles.

It is said that:

... the path of a comet around the Sun has a large eccentricity.

a)	a) Explain the above statemen	nt.		

[1 mark]

b) Figure 3 shows part of the path of a comet close to the Sun.

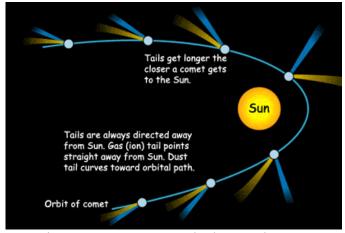


Figure 3. A comet's path close to the Sun.

Briefly explain why the comet's speed is seen to increase as it approaches the Sun.
[2 marks

c) Briefly	explain why the comet's tail becomes more extended when it is closer to the s	Sun.
		[2 marks
		[2 marks
d) There a these two t	are two types of comet, called <i>long-period</i> and <i>short-period</i> . Explain the diffe types.	rence between
		[2 marks]
the outer li	Period comets are thought to be predominantly from the Oort Cloud. This cloud limit of the solar system. Comets are periodically sent from this cloud towards ow this periodic influence on the Oort cloud may be occurring.	
		[2 marks
	explain why it is thought that comets may have played an important role in the planet Earth.	-
		[1 mark]

3. This question is about the different types of galaxy which are observed in the universe.

The Hubble classification of galactic types described four main types of galaxy:

- Ellipticals (E0 to E7).
 Spirals (barred and unbarred) (Sa to Sc / SBa to SBc).
- > Lenticulars.
- > Irregulars.

Explain what the designations 'E0' to 'E7' mean.	
	. .
	. .
[2 mark	CS
There are two types of spiral galaxy – barred and non-barred. State what causes the observed bar and it indicates about the galaxy.	d
[2 mark	KS
Briefly state the similarities and differences between spiral and elliptical galaxies.	
	. •
	•
	. •
[3 mark	KS

d) Classify the galaxies shown in Figure 4 as precisely as you feel you can. In each case, give a single specific type of galaxy.

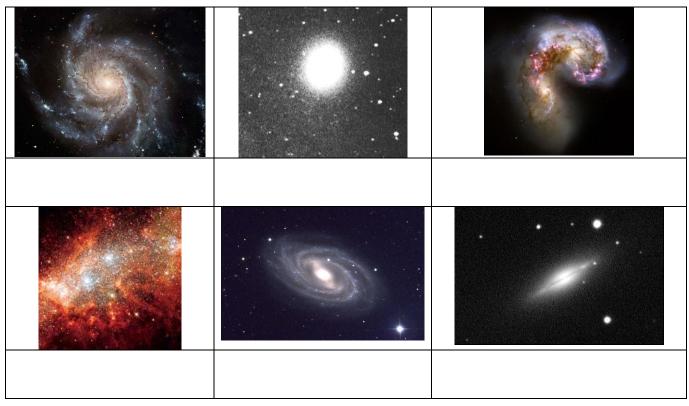


Figure 4. Examples of different types of galaxy.

[3 marks]

[2 marks]

4. This question is about the shape of space-time.

In our present understanding of space-time, the local shape of space-time is distorted by objects with mass. Overall, however, the universe can have one of three possible shapes:

- > Flat space-time.
- > Spherical space-time.
- > Saddle-shaped space-time.

These three shapes are shown in Figure 5.

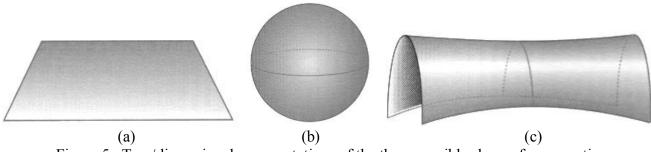


Figure 5. Two/dimensional representations of the three possible shapes for space-time.

a) For each of the tests above, explain how they are able to distinguish between the three possible shapes

Three tests that could be performed to distinguished between these shapes are:

- > Considering parallel straight lines.
- > The internal angles of a triangle.
- The circumference of a circle.

for space-time	:				
Parallel utraigh	nt lines :	 	 		
				[2	2 marks]
Internal angles	of a triangle:	 	 		
		 	 		• • • • • • • • • • • • • • • • • • • •
		 	 		2 marks]
Circumference	of a circle:	 	 		

For each possi critical density	-	time shape, s	state the cor	nditions pla	aced upon	the density	of space co	mpared with
Flat space-tim	ie	:						
Spherical space	ce-time	:						
Saddle/shaped	l space-tim	ne:						
								[2 marks]
c) For a spher universe.	rical space-	-time, briefly	explain wł	nat this pre	dicts for th	e continue	d expansion	of the
							• • • • • • • • • • • • • • • • • • • •	
							• • • • • • • • • • • • • • • • • • • •	
							• • • • • • • • • • • • • • • • • • • •	[2 marks]

b) The parameter which can be said to distinguish between these three options is the density of space (ρ)

compared with what is known as critical density (ρ_o).

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Section B

Answer ONE question only.

You are advised to read BOTH questions before starting your answer and spend no more than 20 minutes on this section

Black Hole Eclipse

NASA's Chandra X-ray Observatory has observed a remarkable eclipse of a supermassive black hole, allowing a disk of hot matter swirling around the hole to be measured for the first time.

The supermassive black hole is located in NGC 1365, a spiral galaxy 60 million light years from Earth. It contains a so-called active galactic nucleus, or AGN. Scientists believe that a black hole at the centre of the AGN is fed a steady stream of material from a surrounding disk. Matter just about to fall into a black hole should be heated to millions of degrees before passing over the event horizon, or point of no return. This super-heated disk material glows brightly in the X-ray part of the electromagnetic spectrum where Chandra can see it.



Figure 6. The black hole eclipse occurred in the core of this barred spiral galaxy, NGC 1365.

The disk of gas around the central black hole in NGC 1365 is much too small to resolve directly with a telescope. However, the disk was eclipsed by an intervening cloud. By recording the time taken for the disk to go in and out of eclipse, scientists were able to estimate the diameter of the disk.

The Chandra team directly measured the diameter of the X-ray source as about seven times the distance between the Sun and the Earth or 7 AU. For comparison, if such a disk were placed in our own solar system, it would swallow all the planets out to Mars and most of the asteroid belt as well.

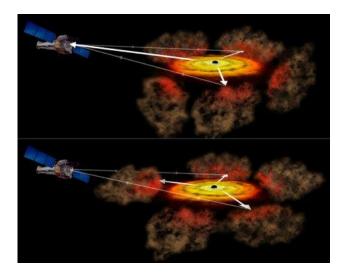


Figure 7. An artist's concept of Chandra observing the black hole eclipse.

According to these measurements, the source of X-rays is about 2 billion times smaller than the host galaxy NGC 1365 and only about 10 times larger than the estimated size of the black hole's event horizon. This is consistent with theoretical predictions.

"Thanks to this eclipse, we were able to probe much closer to the edge of this black hole than anyone has been able to before," said co-author Martin Elvis from CfA. "Material this close in will likely cross the event horizon and disappear from the universe in about a hundred years, a blink of an eye in cosmic terms."

In addition to measuring the size of this disk of material, Risaliti and his colleagues were also able to estimate the location of the dense gas cloud that eclipsed the X-ray source and central black hole. The Chandra data show that this cloud is one hundredth of a light year from the black hole's event horizon-much closer than anyone expected. So this is a bit of a puzzle.

"AGN [are among] the brightest objects in the cosmos and they are powerful probes of the early history of the Universe. It's vital we understand their basic structure," said Risaliti. "It turns out that we still have work to do to understand these monsters."

Black Hole Eclipse

a) Explain the	following terms used in the article:	
Black Hole:		
Spiral Galaxy:		
Event Horizon	:	
		[3 marks]
b) How far awa	ay is NGC1365 from the Earth? Give your answer in metres.	
		[2 marks]
c) The article s characteristics of	speculates that NGC1365 contains an active galactic nucleus (AGN). Give TWO of an AGN.	
Characteristic 1	l:	
Characteristic 2	2:	
		 [2 marks]

d) Why is su	ch a high tempe	rature needed in orde	er to generate x		M11/4/ASTRO/SP2/ENG/XX
a) willy is sur					
					[1 mark]
		GN will almost cert ogen and explain wh			ONE box below to
		Molecular			
		Atomic			
		Plasma			
		Tick (ONE box only	_	
					[2 marks
	-	lone, if the temperat		_	
2,000,000K, d	do you think nuc	elear fusion is occurr	ring in the gasA	Explain your a	nswer.
					[2 marks]
	eter of the event eely 3.5 x 10 ³⁷ kg		0.7AU. Use thi	s data to show t	hat the mass of the AGN
					[3 marks]

h) If NGC1365 is similar to the Milky Way (Mass approximately $10^{11} \rm M_{\odot}$) what percentage of its mass is contained within the central black hole?
[2 marks]
The article refers to observations of galaxies such as NGC1365 as probes to the early history of the universe.
When we see NGC1365 now, how long has it taken for this light to reach us? Give your answer in years.
[1 mark]
i) If another civilisation were viewing the same galaxy but from only half the distance, how would you expect their view of the emitted radiation to differ from ours?
[2 marks]
** END OF OUESTION **

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Great Perseids

Got a calendar? Circle this date: Sunday, August 12th. Next to the circle write "all night" and "Meteors!" Attach the above to your refrigerator in plain view so you won't miss the 2007 Perseid meteor shower.

"It's going to be a great show," says Bill Cooke of NASA's Meteoroid Environment Office at the Marshall Space Flight Centre. "The Moon is new on August 12th--which means no moonlight, dark skies and plenty of meteors." How many? Cooke estimates one or two Perseids per minute at the shower's peak.



Figure 8. A Perseid fireball photographed August 12, 2006, by Pierre Martin of Arnprior, Ontario, Canada.

The source of the shower is Comet Swift-Tuttle. Although the comet is nowhere near Earth, the comet's tail does intersect Earth's orbit. We glide through it every year in August. Tiny bits of comet dust hit Earth's atmosphere travelling 132,000 mph. At that speed, even a smidgen of dust makes a vivid streak of light--a meteor--when it disintegrates. Because Swift-Tuttle's meteors fly out of the constellation Perseus, they are called "Perseids."

"Earthgrazers are long, slow and colourful; they are among the most beautiful of meteors," says Cooke. He cautions that an hour of watching may net only a few of these--"at most"--but seeing even one makes the long night worthwhile.

As the night unfolds, Perseus climbs higher and the meteor rate will increase many-fold. "By 2 am on Monday morning, August 13th, dozens of Perseids may be flitting across the sky every hour." The crescendo comes before dawn when rates could exceed a meteor a minute.

For maximum effect, Cooke advises, "get away from city lights." The brightest Perseids *can* be seen from cities, he allows, but the greater flurry of faint, delicate meteors is visible only from the countryside. Scouts, this is a good time to go camping.

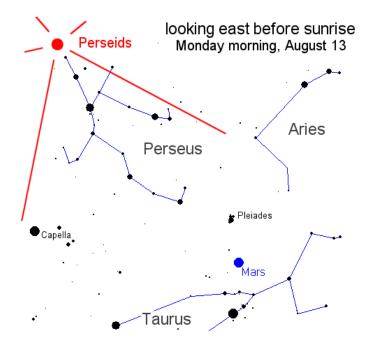


Figure 9. The eastern sky, viewed during the hours before sunrise on Monday, Aug. 13, 2007.

And there's a bonus: Mars. In the constellation Taurus, just below Perseus, Mars shines like a bright red star. Many of the Perseids you see on August 12th and 13th will flit right past it. Instead of following the meteor, you may find you have a hard time taking your eyes off Mars. There's something bewitching about it, maybe the red colour or perhaps the fact that it doesn't twinkle like a true star. You stare at Mars and it stares right back.

Earth and Mars are converging for a close encounter in December 2007. NASA is taking advantage by launching a new mission to Mars--the Phoenix Lander. Phoenix will touch down on an arctic plain where it can dig into the ground and investigate layers of soil and ice, searching for, among other things, a habitable zone for primitive microbes. The launch window opens on August 3rd, so by the time the Perseids arrive Phoenix may be hurtling toward the Red Planet. Landing: late Spring 2008.

Great Perseids

a) Explain the	e following terms used in the article:	
Comet:		
Constellations	s:	
		••••
	[2 m	 arks]
b) State one s	similarity and one difference between a New Moon and a Lunar Eclipse.	
Similarity:		
Difference:		••••
	[2 m	arks]
· -	e fact that the comet Swift-Tuttle has long since left the Earth's vicinity, explain why we Perseid meteor shower every August.	e
	[1 n	nark]
d) When parti how this light	icles of dust enter the Earth's atmosphere, fireballs can occur as shown in Figure 8. Expis produced.	plain
	[2 m	 arks]

e) Why are su	ch fireballs rarely seen in the daytime?	
		[1 marks]
particle is that	dust particle entering the Earth's atmosphere as discussed in (d). One possible fait grazes the atmosphere, entering the edge and then leaving again, back into spawo other possible fates for such a particle.	
Possible fate:		
Possible fate :		
		[2 marks]
g) The article so.	refers to the fact that the shower is best viewed away from city lights. Explain v	why this is
		[2 marks]
h) Briefly disc	cuss a"different reason why the Pleiades shower might not be seen in the night sk	y.
i) Why does N	Mars look red in the night sky?	[2 marks]
		[1 mark]

Explain this motion and state if such motion is also observed for the planet Venus. [3 marks] k) The final paragraph refers to the proximity of Earth and O ars as an opportunity to investigate Mars with the Phoenix Lander. Despite the fact that this would be very costly, give TWO 'r quuldrg'dgpghku' qh'uwej "cp"gzr mtc wqp" Dgpghkv1: Dgpghkv2: [2 marks]

j) The article comments that Mars looks different to a star in the night sky because it does not twinkle. However, when viewed over a period of many weeks, Mars could be seen to show retrograde motion

** END OF QUESTION **

against the background of stars.

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